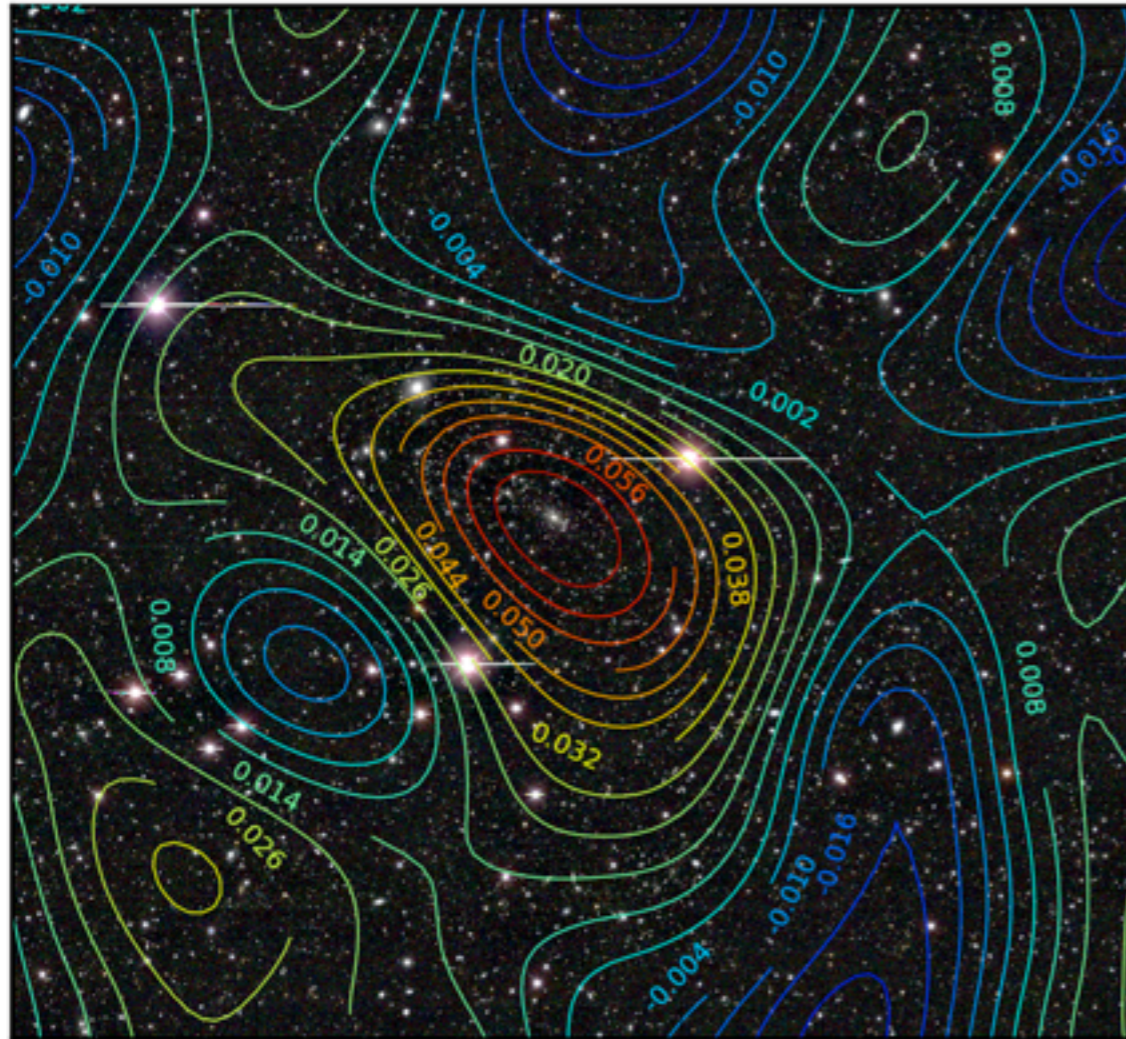


# The BIGGER picture

Mass & light maps of massive galaxy clusters from early DECam data



Bob Armstrong	Elisabeth Krause
David Bacon	Huan Lin
Keith Bechtol	Niall McCrann
Sarah Bridle	Ken Patton
Fabrice Brimiouille	Andres Plazas
Joseph Clampitt	Barnaby Rowe
Daniel Gruen	Eduardo Rozo
Michael Hirsch	Eli Rykoff
Klaus Honscheid	Stella Seitz
Eric Huff	Eric Suchyta
Bhuv Jain	Rafal Szepletowski
Mike Jarvis	Vinu Vikram
Stephanie Jouvel	Julia Young
Tomasz Kacprzak	... and the DES SV squad

Peter Melchior



# Dark Energy Survey

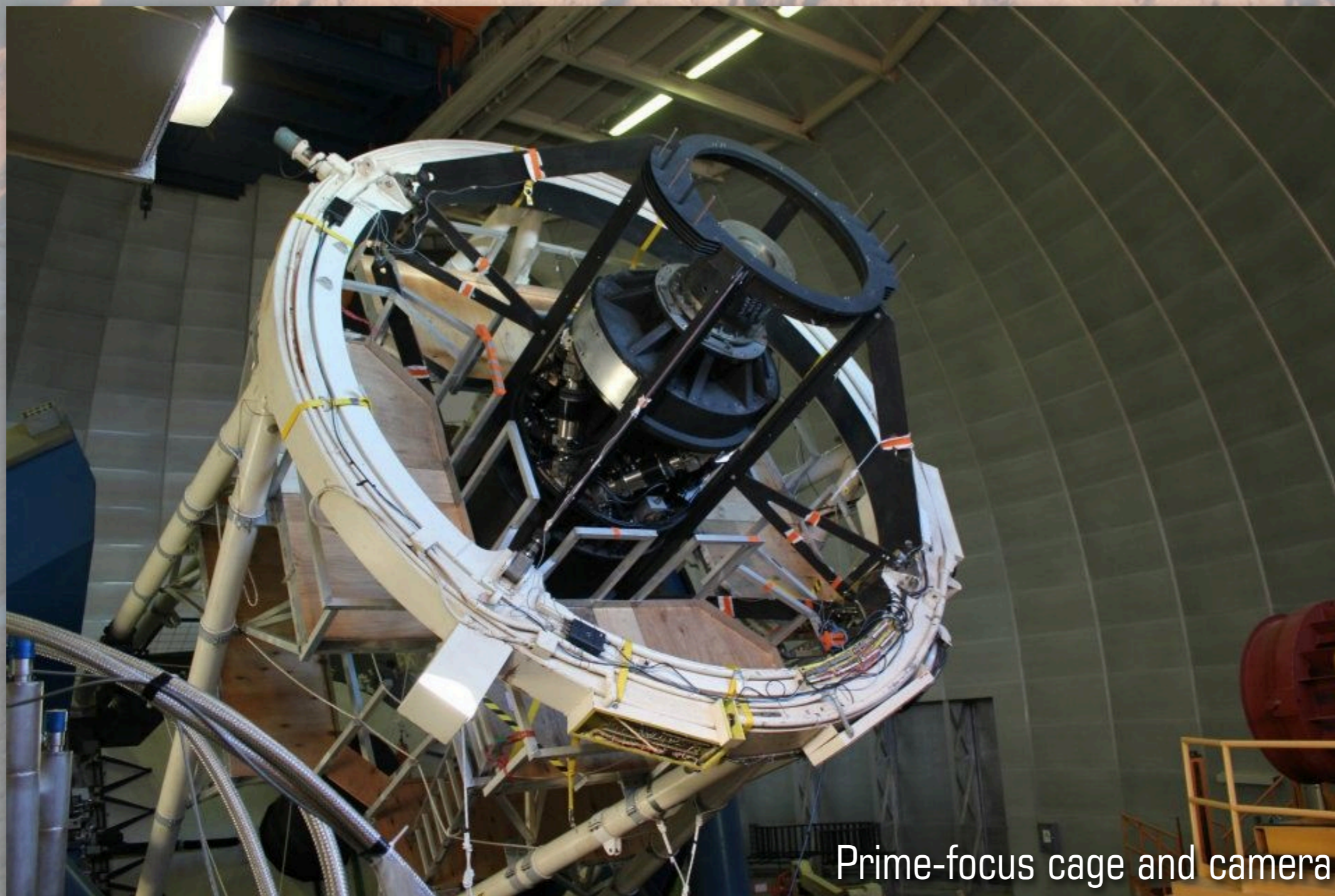


Cerro Tololo Inter-American Observatory

# Dark Energy Survey



DARK ENERGY  
SURVEY

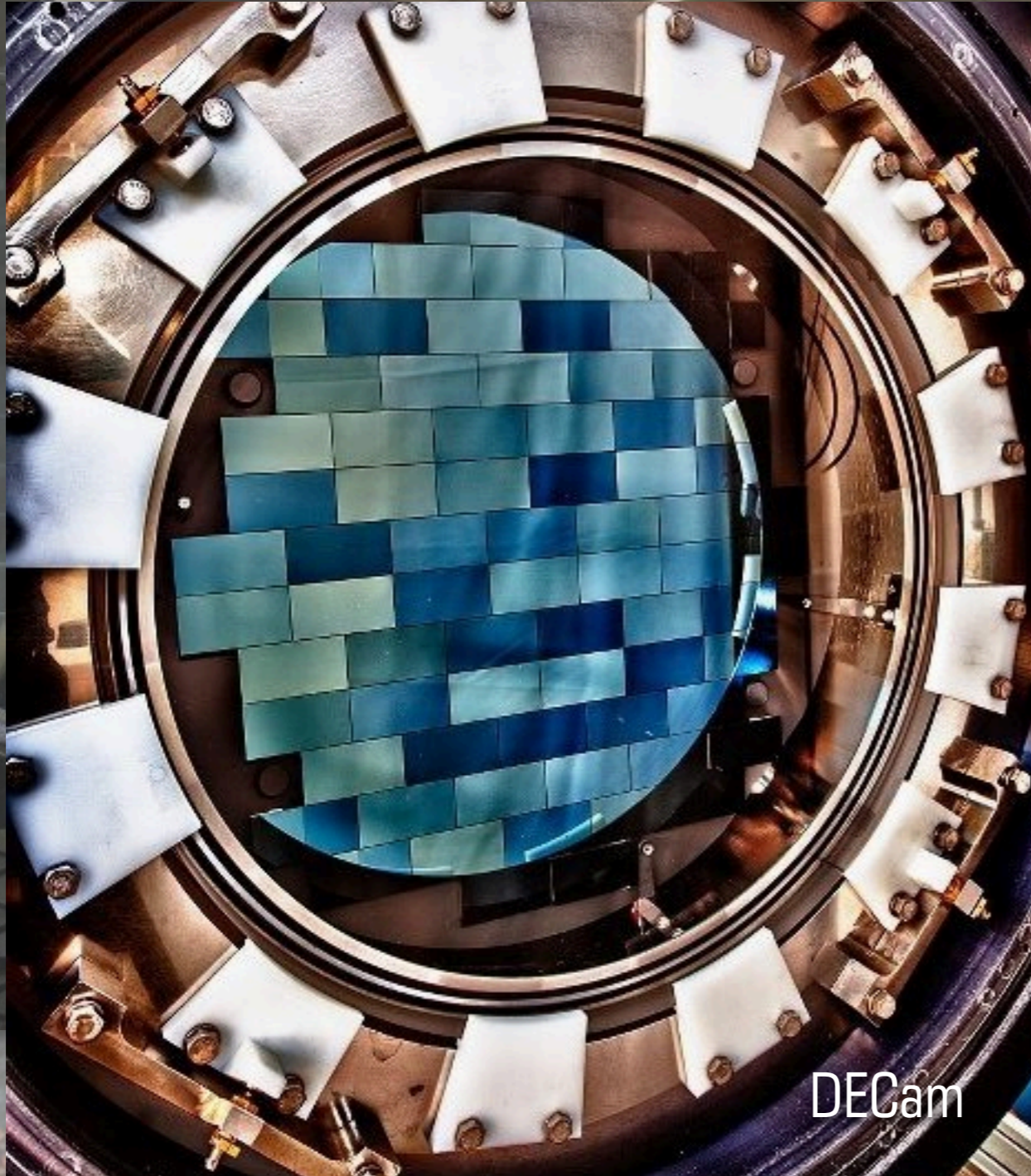


Prime-focus cage and camera

# Dark Energy Survey



DARK ENERGY  
SURVEY

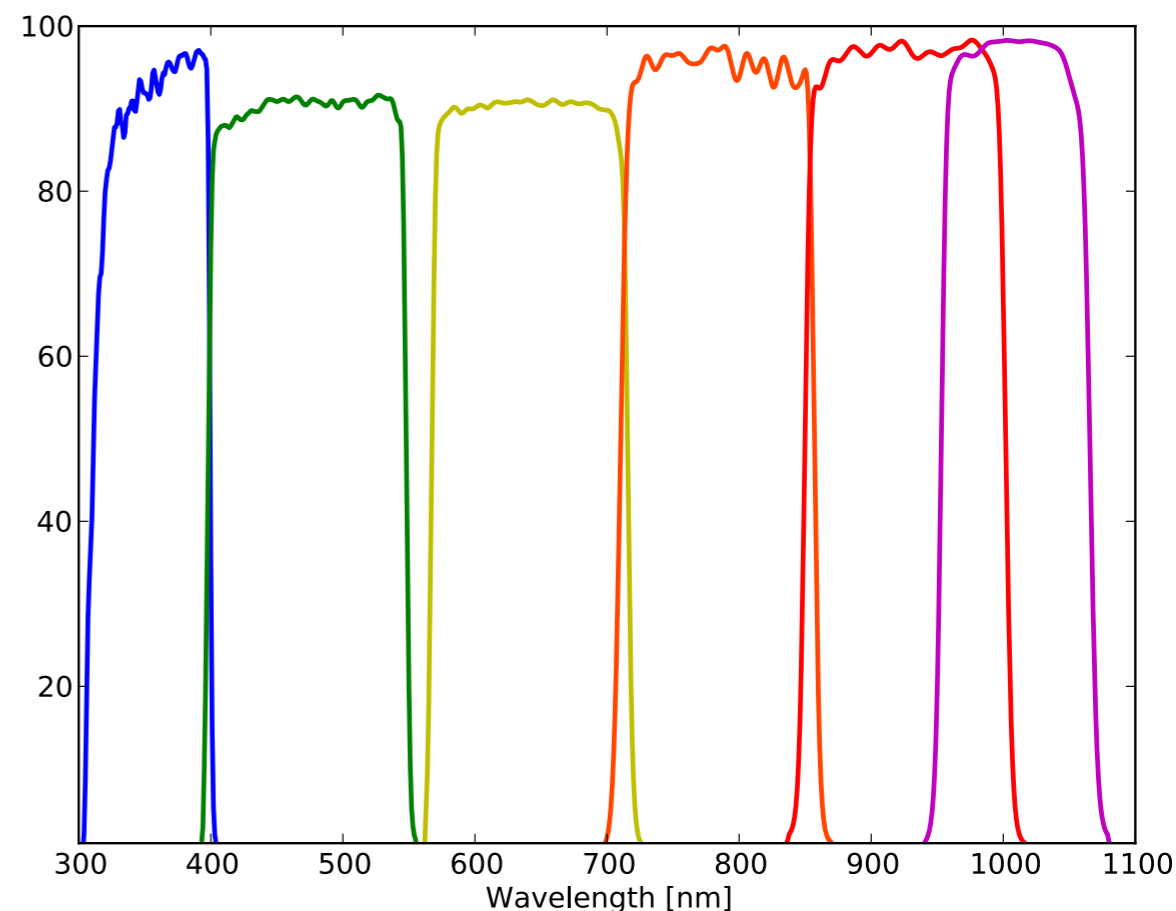
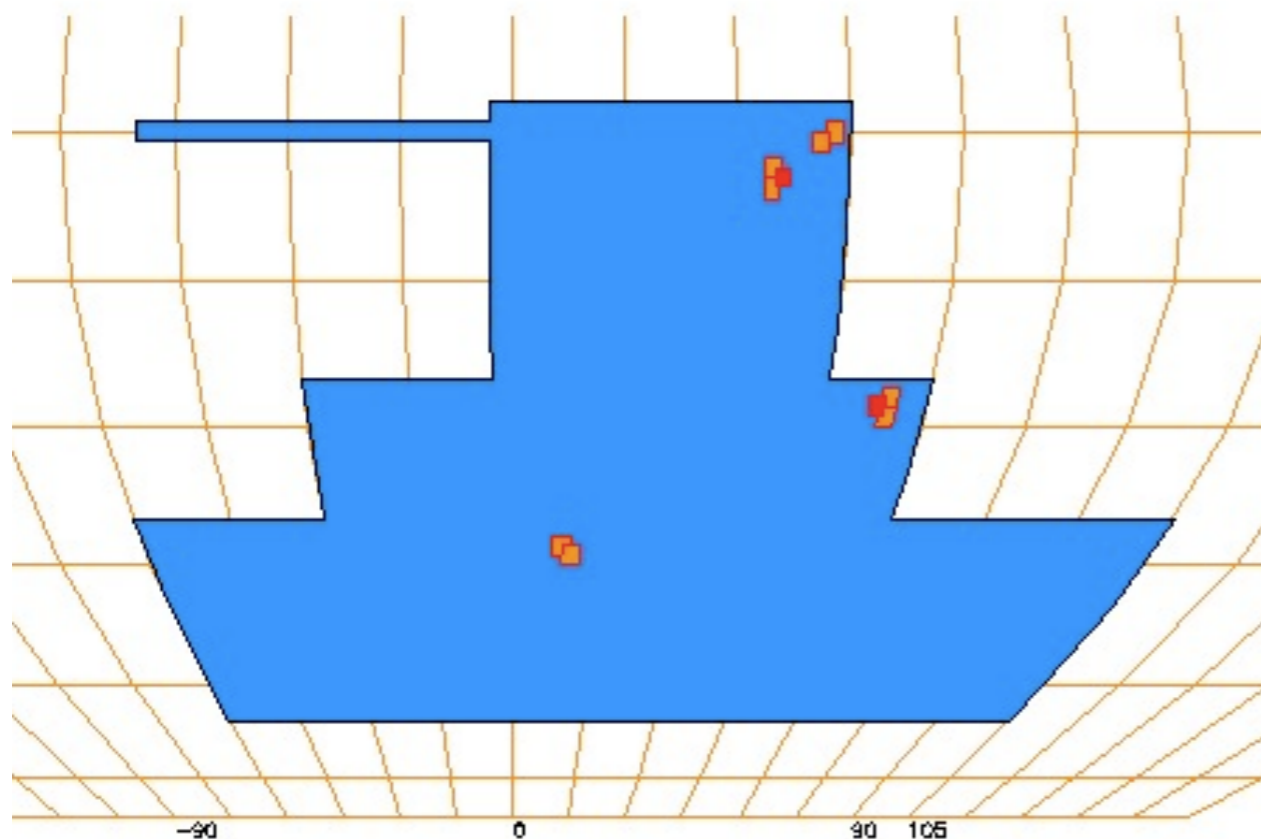


# Dark Energy Survey



g-filter installation

# Survey characteristics



5000 sq. degrees

filters: grizY + U

10 x 90 seconds

limiting magnitudes: 25.2 (g) .. 23.4 (z)

# Clusters in Science Verification



1. validate the data quality delivered by DECam for the purpose of galaxy cluster and lensing studies

2. utilize the large FoV of DECam to create light and mass maps over more than 2 square degrees around these clusters.

# Clusters in Science Verification



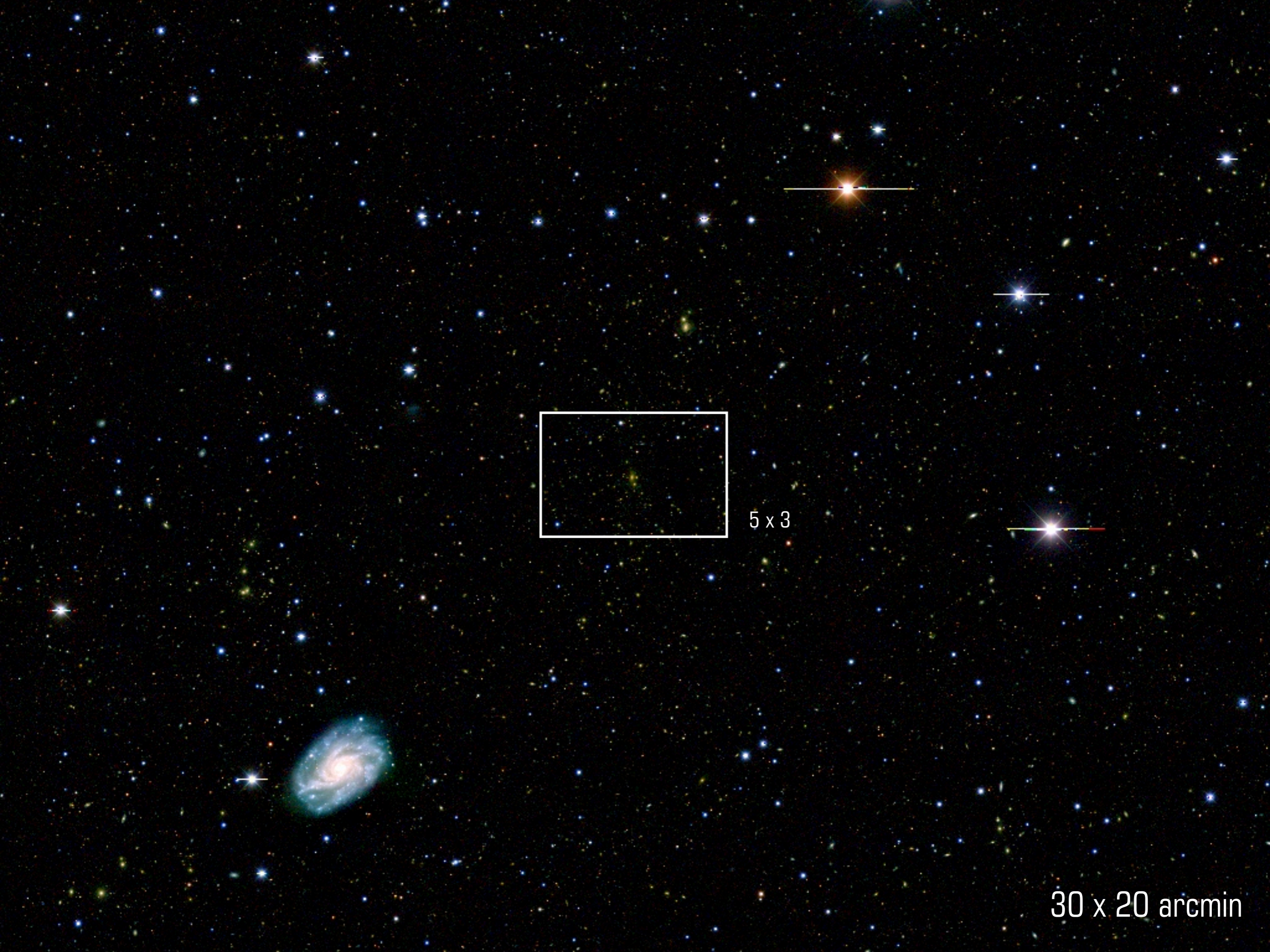
1. validate the data quality delivered by DECam for the purpose of galaxy cluster and lensing studies

2. utilize the large FoV of DECam to create light and mass maps over more than 2 square degrees around these clusters.



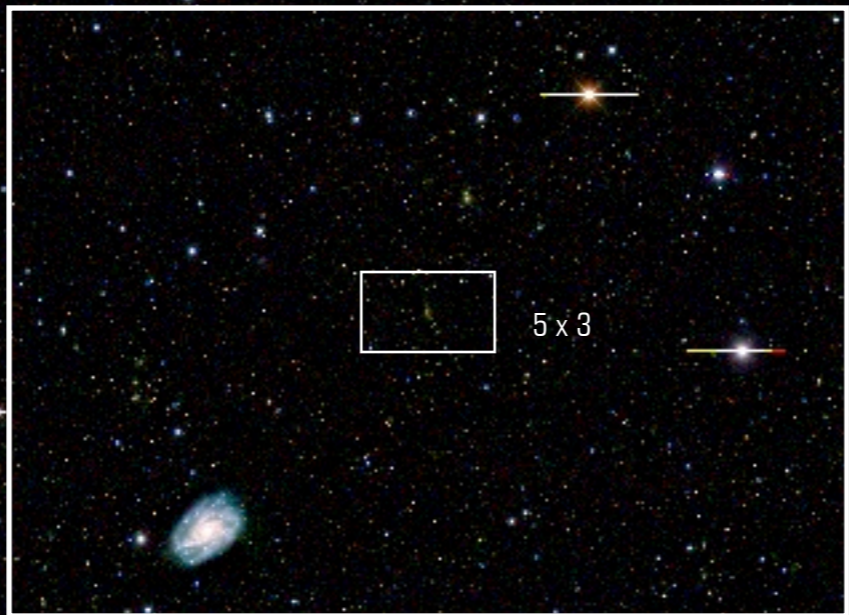


5 x 3 arcmin



5 x 3

30 x 20 arcmin

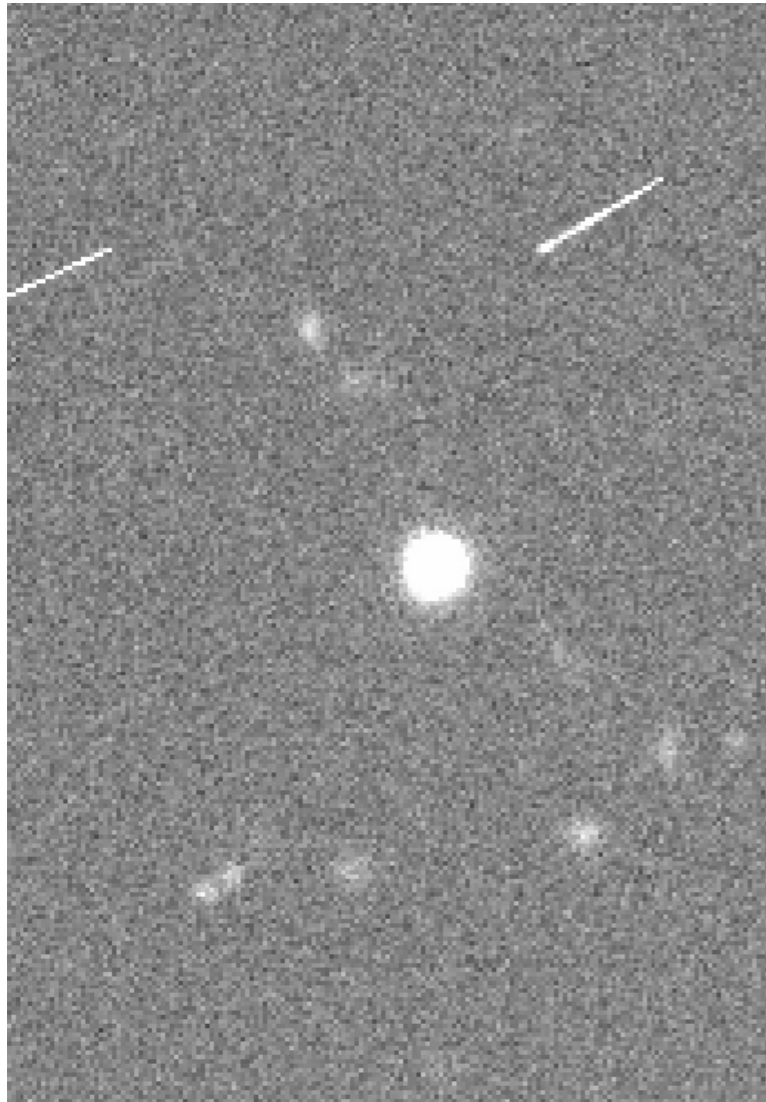


5 x 3

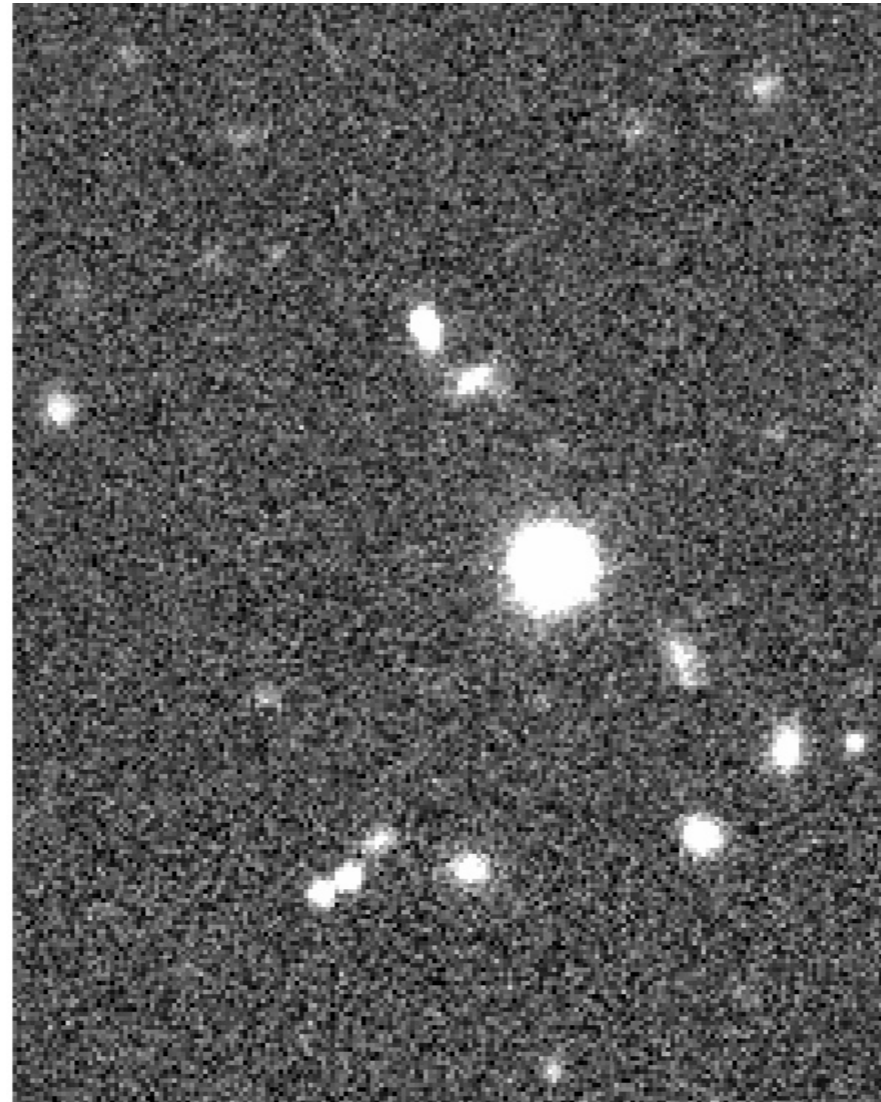
30 x 20

90 x 60 arcmin

# Coadds



single i-band exposure

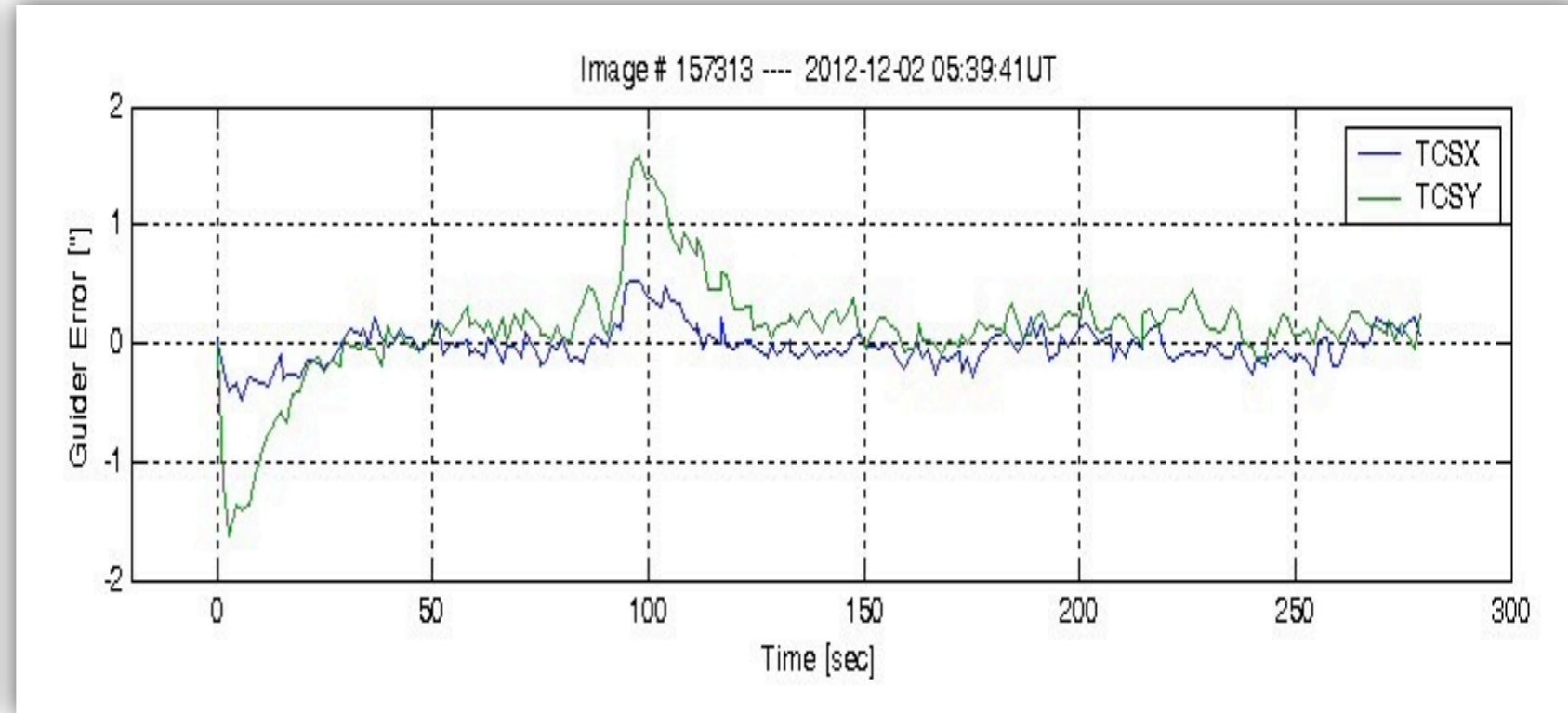
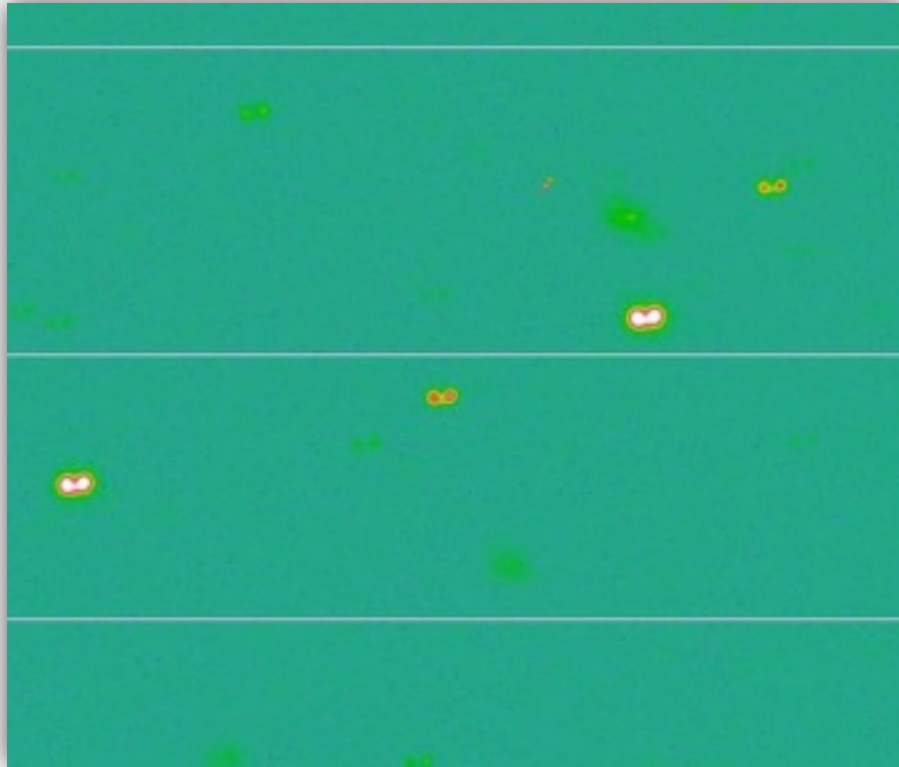


Swarp: median coadd of 10 exp.  
 $n_{\text{gal}} = 21 / \text{sq. arcmin}$



mean coadd+outlier rejection  
 $n_{\text{gal}} = 25 / \text{sq. arcmin}$

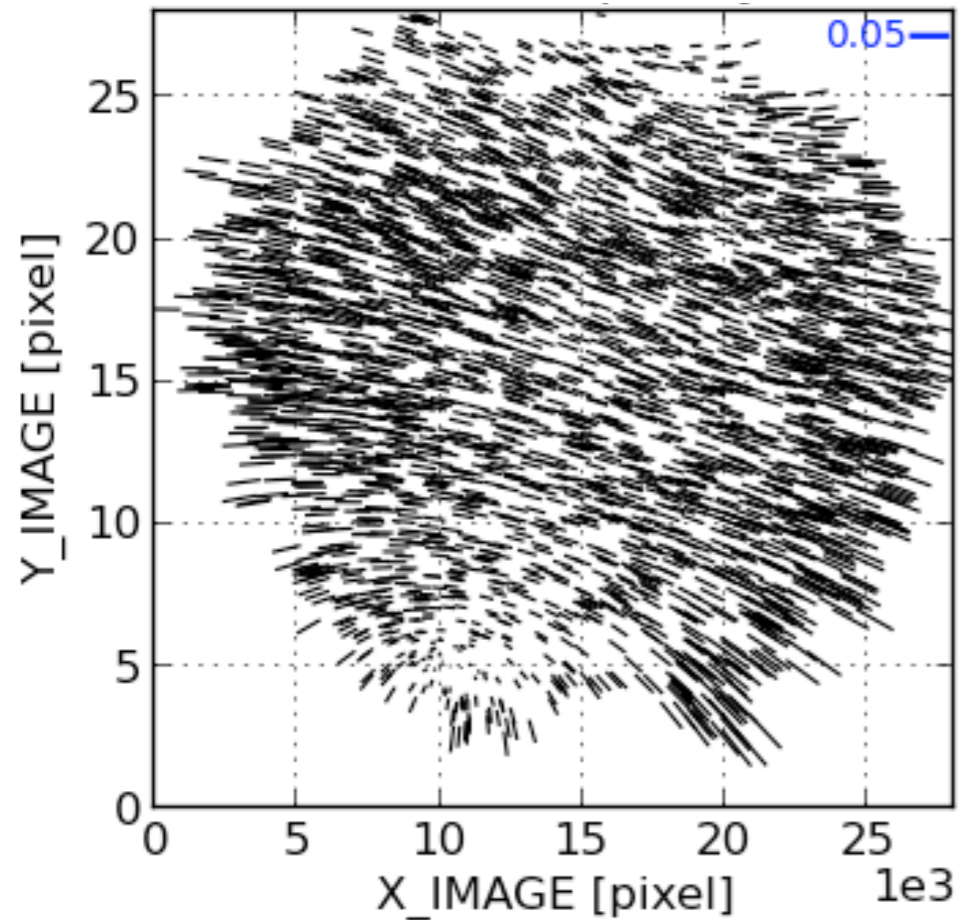
# Early data issues



November 2012:

- ▶ early data affected by “guider jumps”
- ▶ tracking performance not as desired

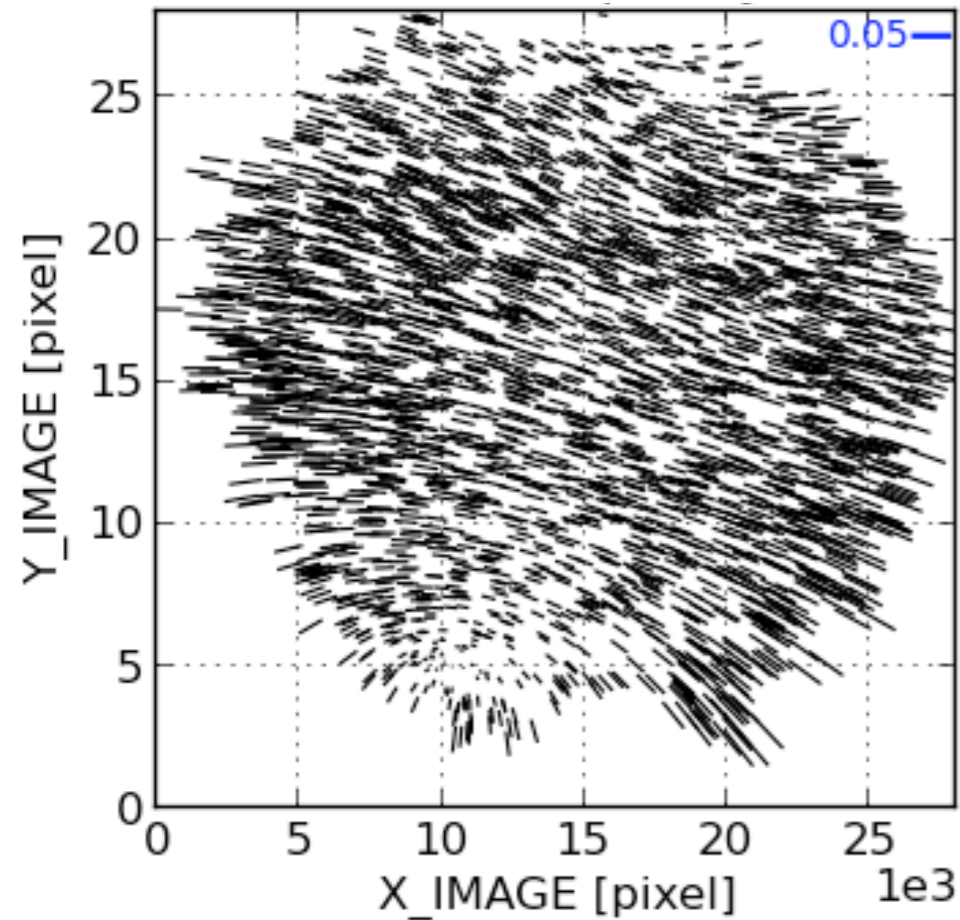
# Early data issues



November 2012:

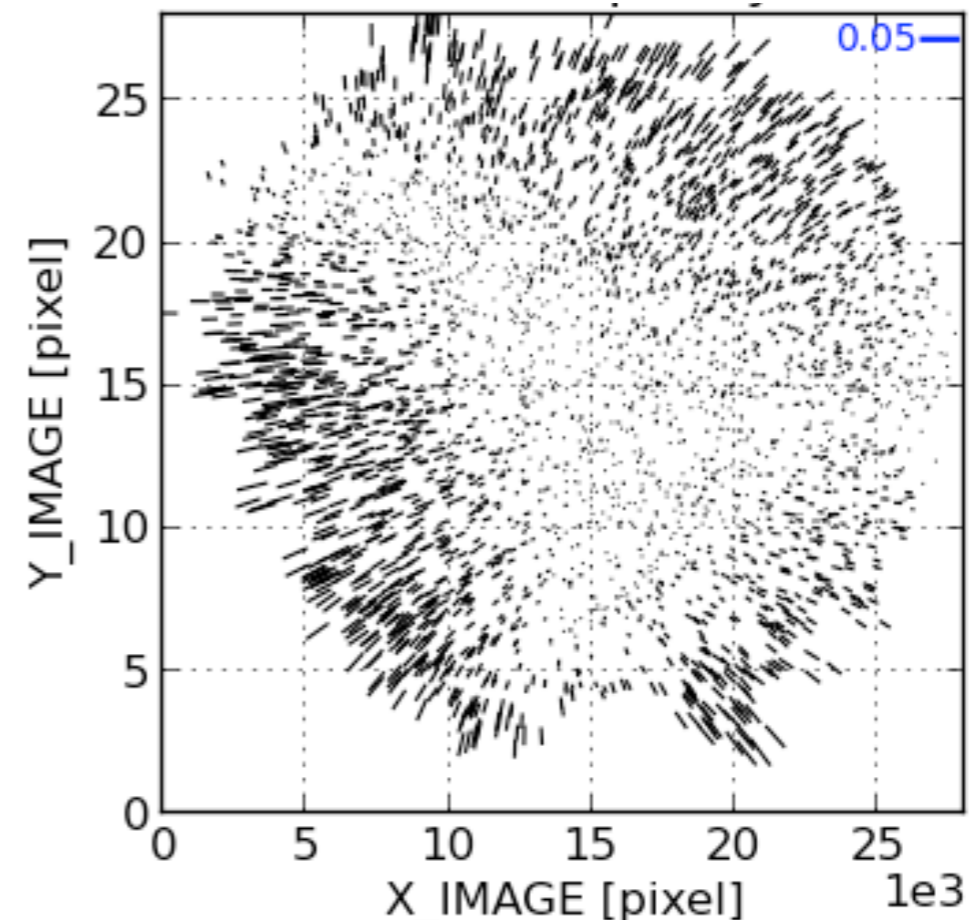
- ▶ early data affected by “guider jumps”
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# Early data issues



November 2012:

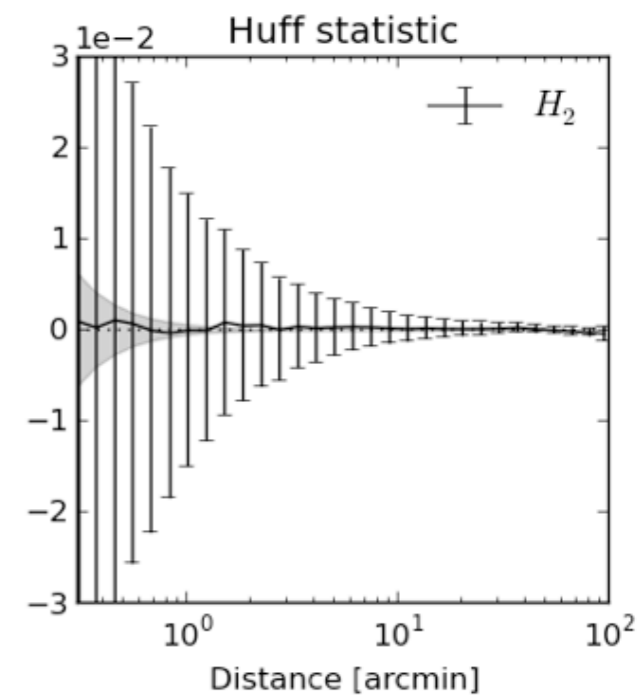
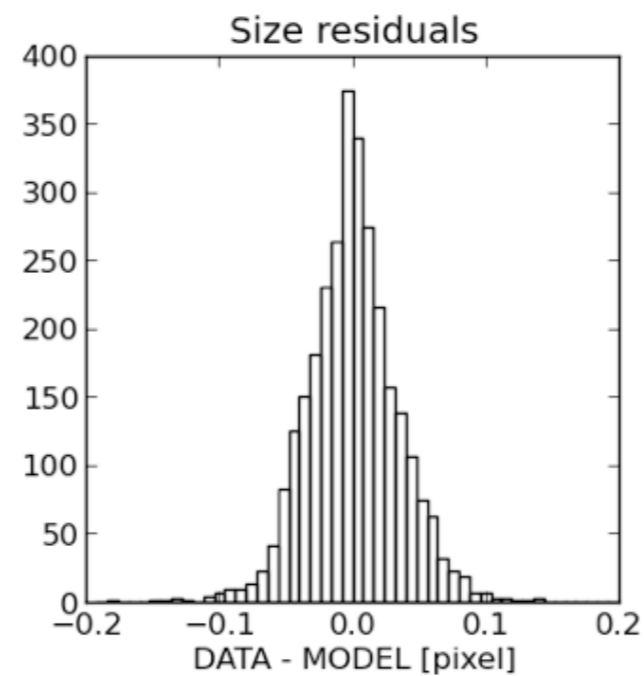
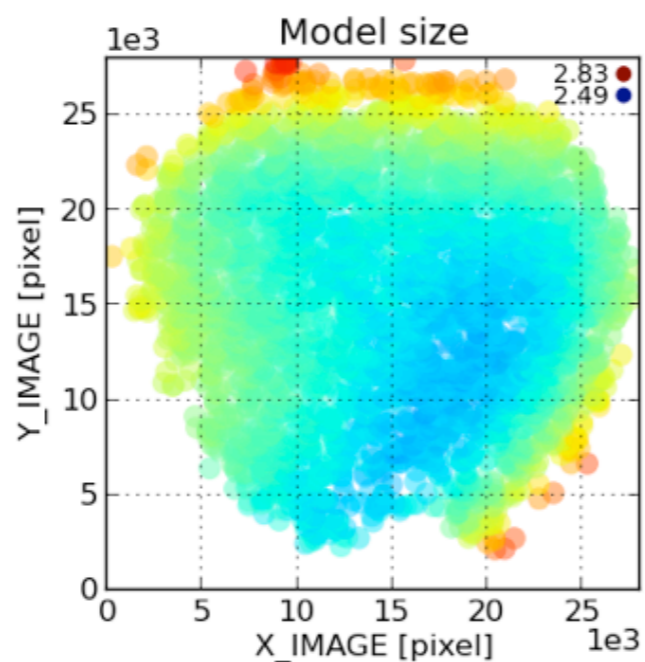
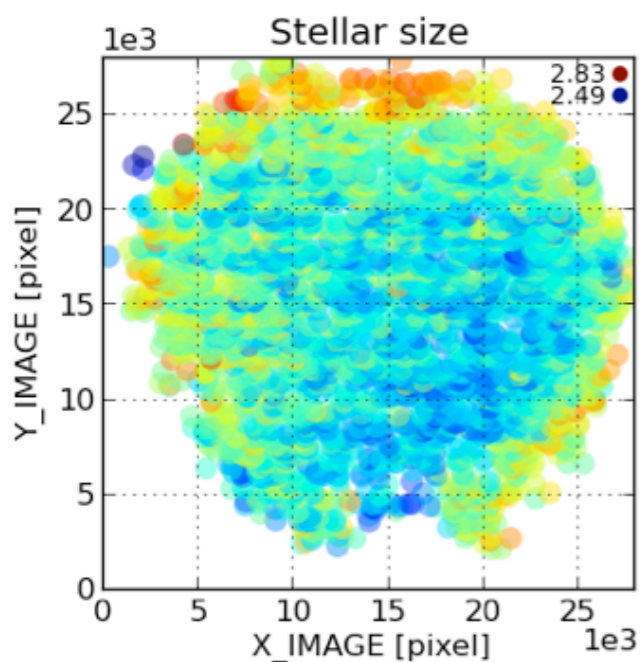
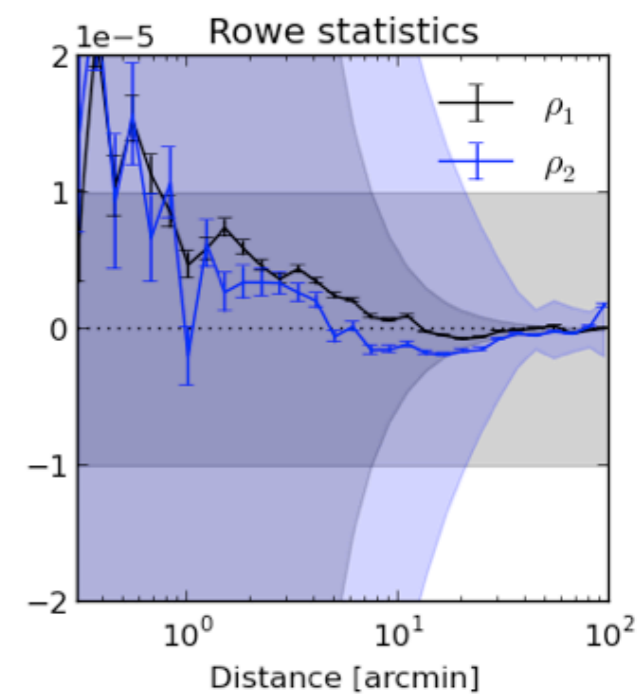
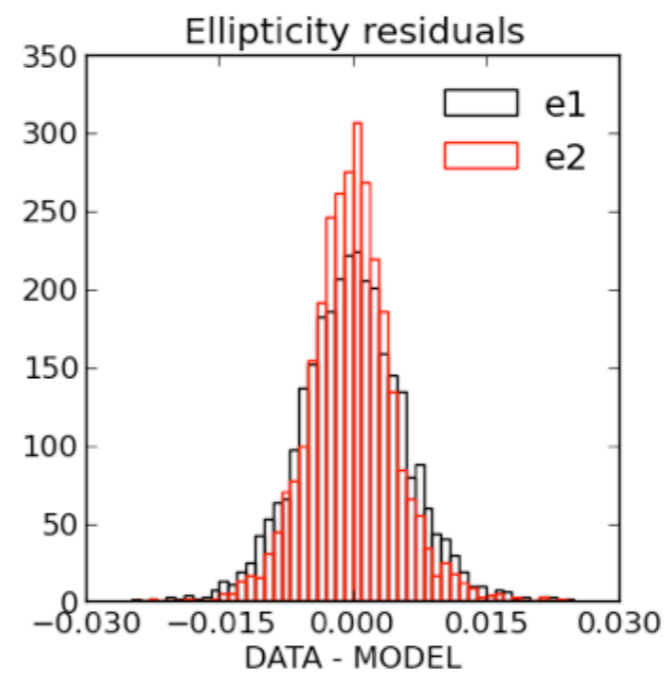
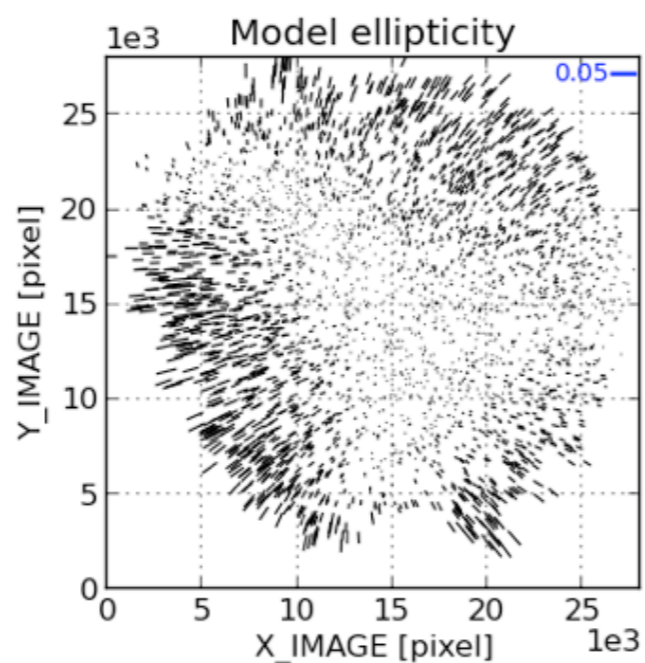
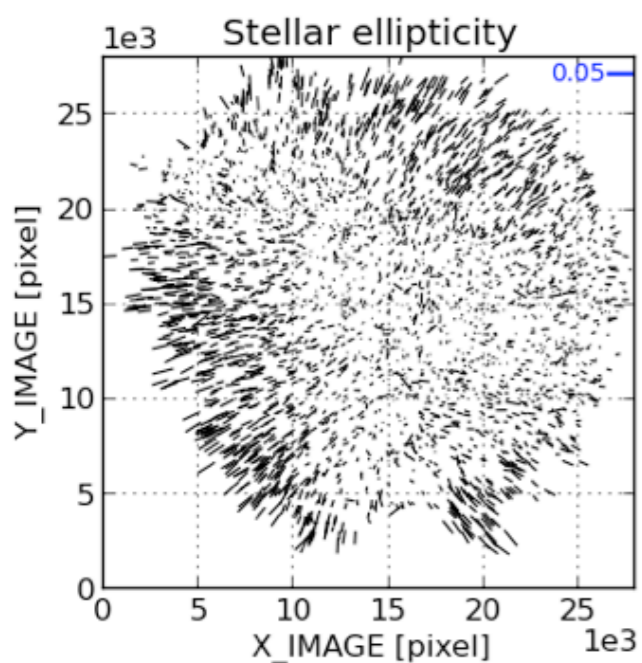
- ▶ early data affected by “guider jumps”
- ▶ tracking performance not as desired



December 2012:

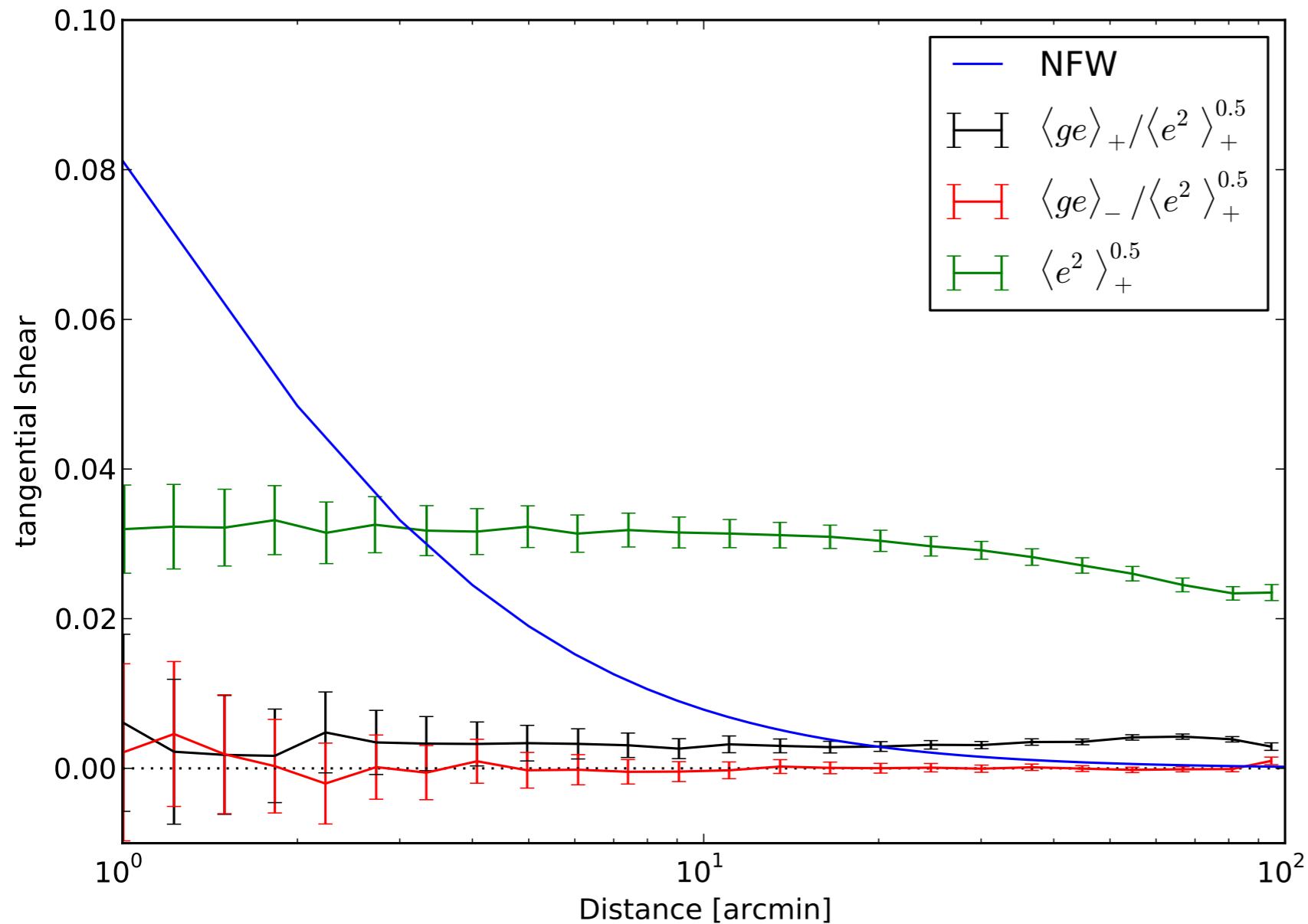
clearly improved image quality

# PSF modeling



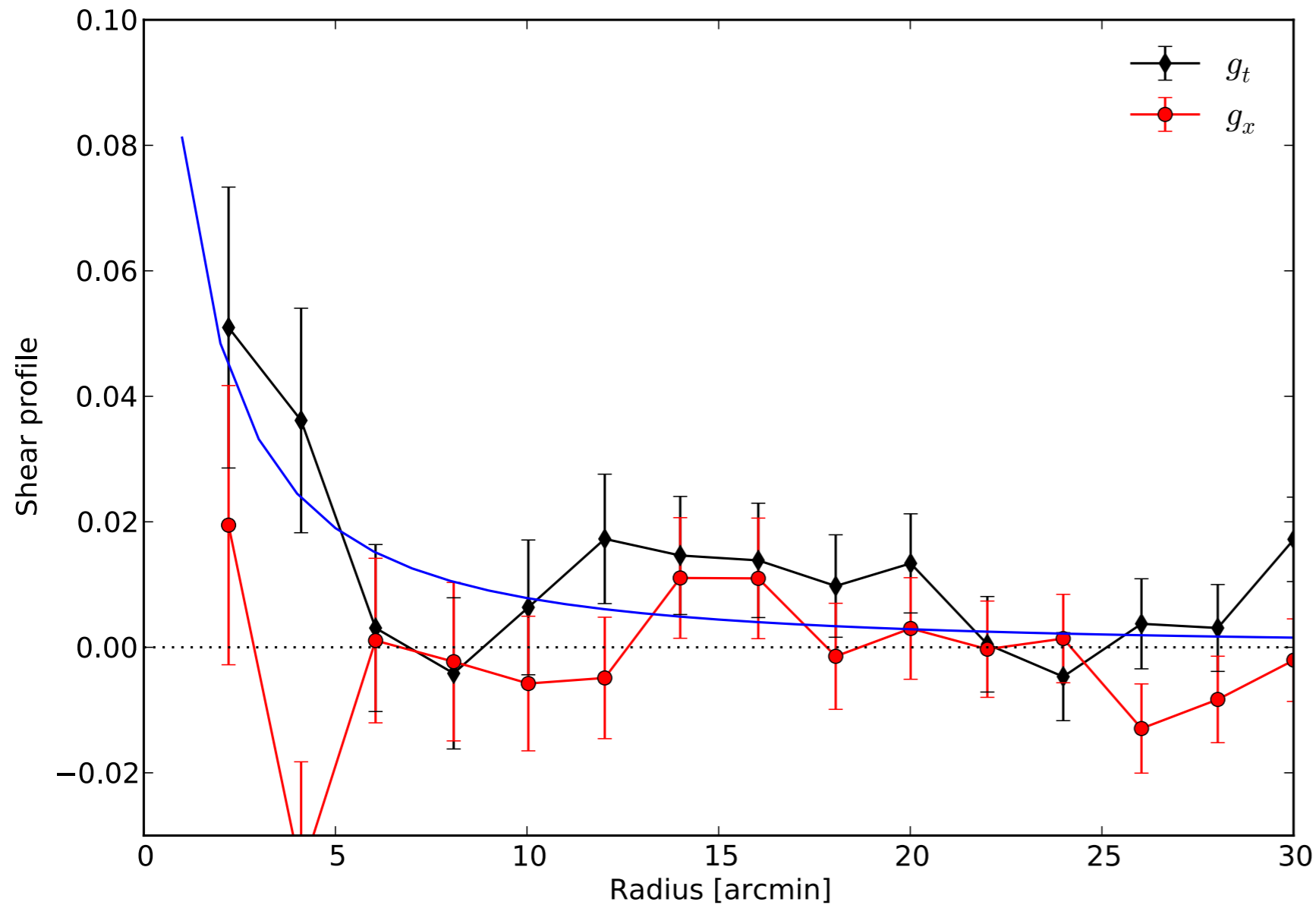


# Star-galaxy correlation function



- ▶ KSB
- ▶ shapelet
- ▶ DEIMOS
- ▶ im3shape

# Shape catalogs



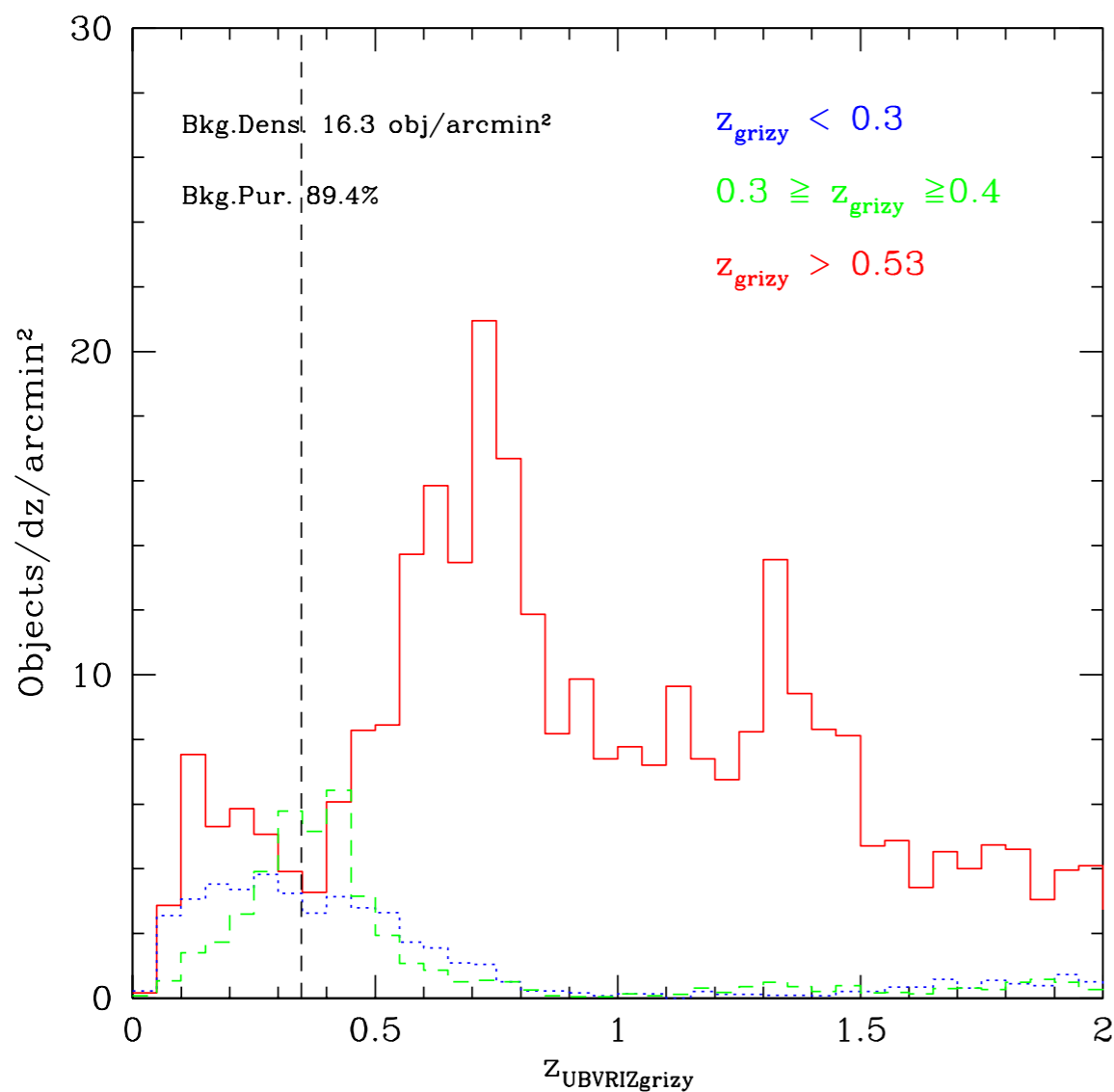
- ▶ KSB
- ▶ shapelet
- ▶ DEIMOS
- ▶ im3shape

$n_{\text{gal}} \sim 15/\text{sq. arcmin}$

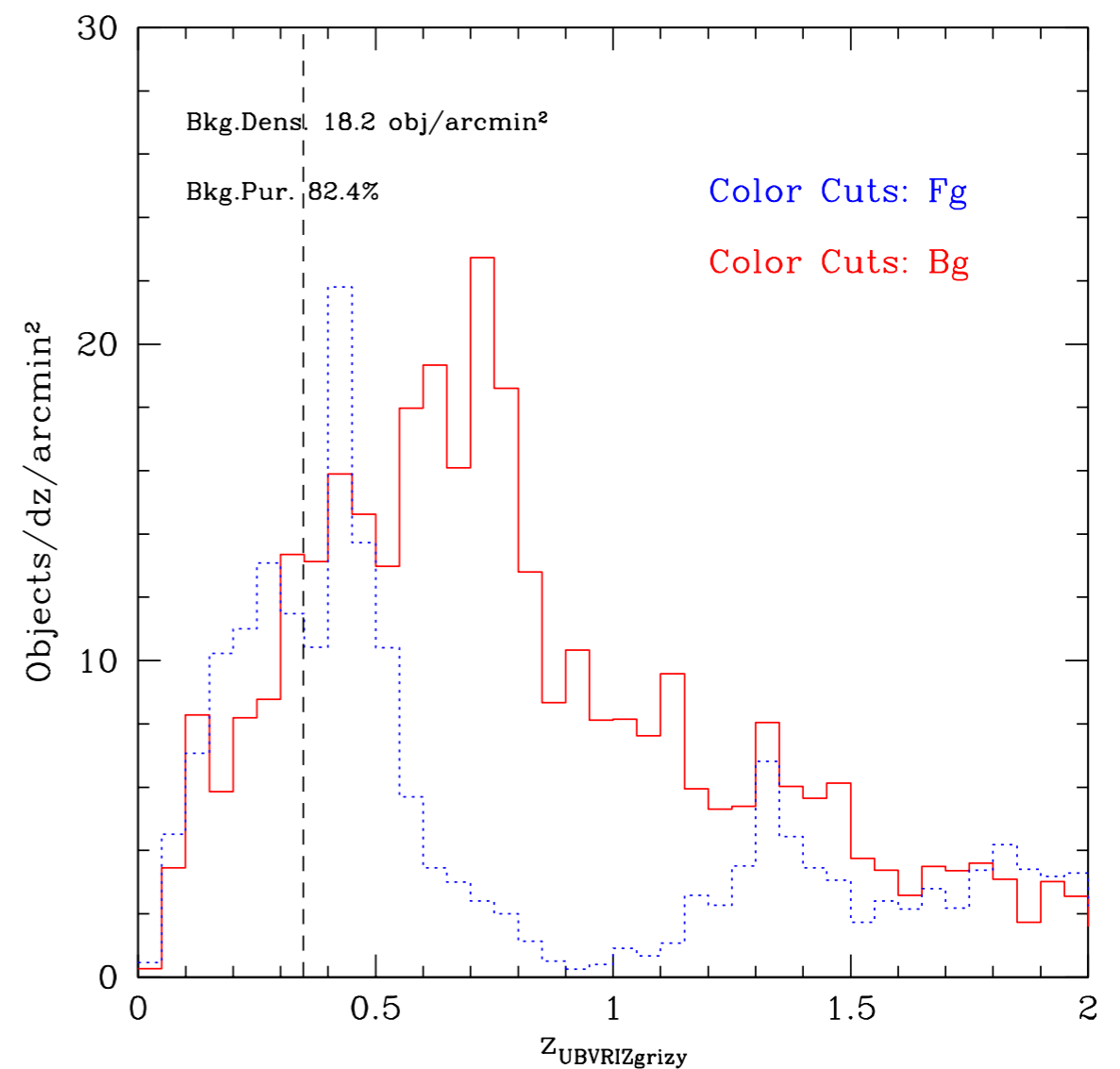
# Photometric catalogs



Foreground / Background selection based on color cuts and photo-zs



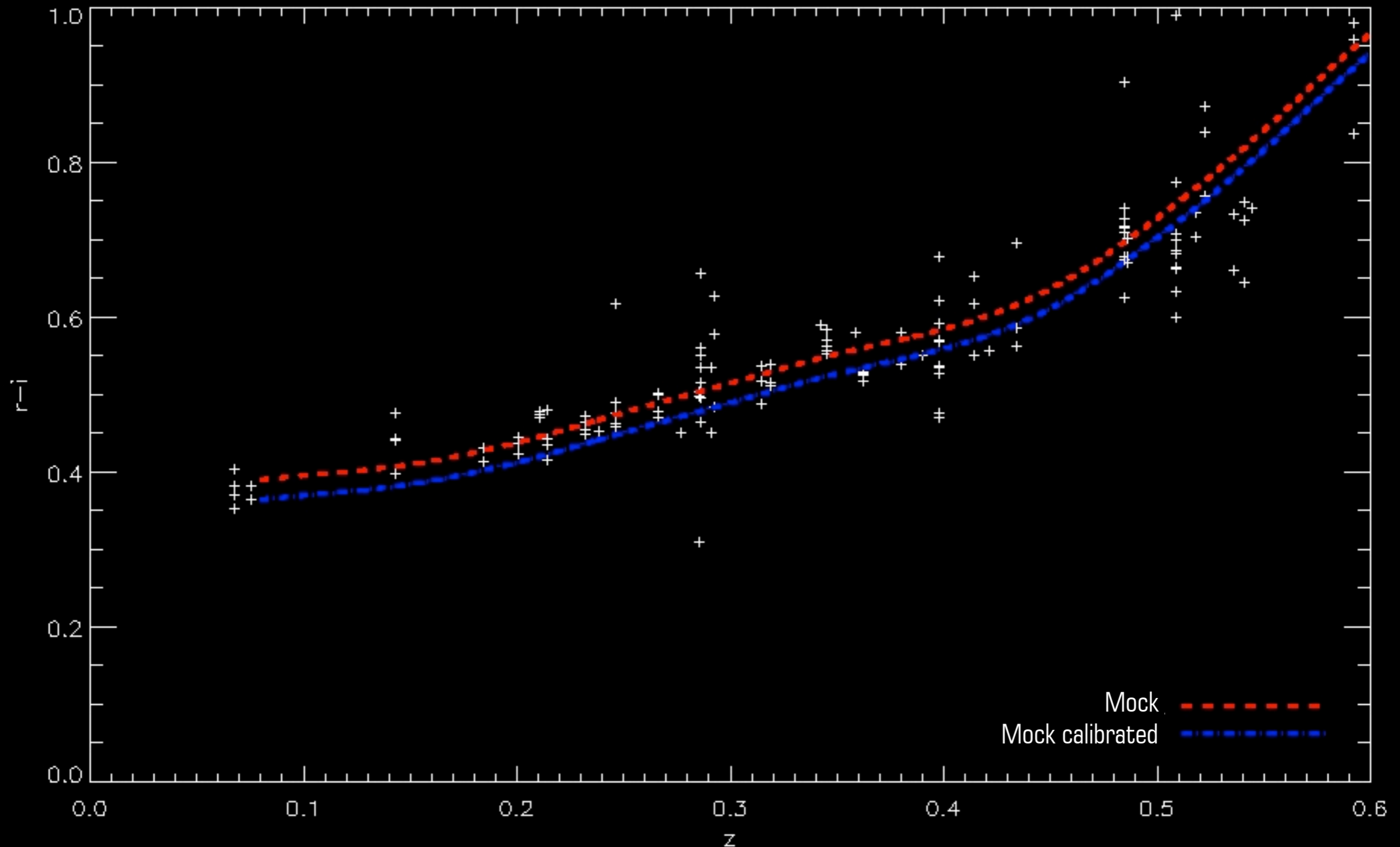
DES photo-z selection



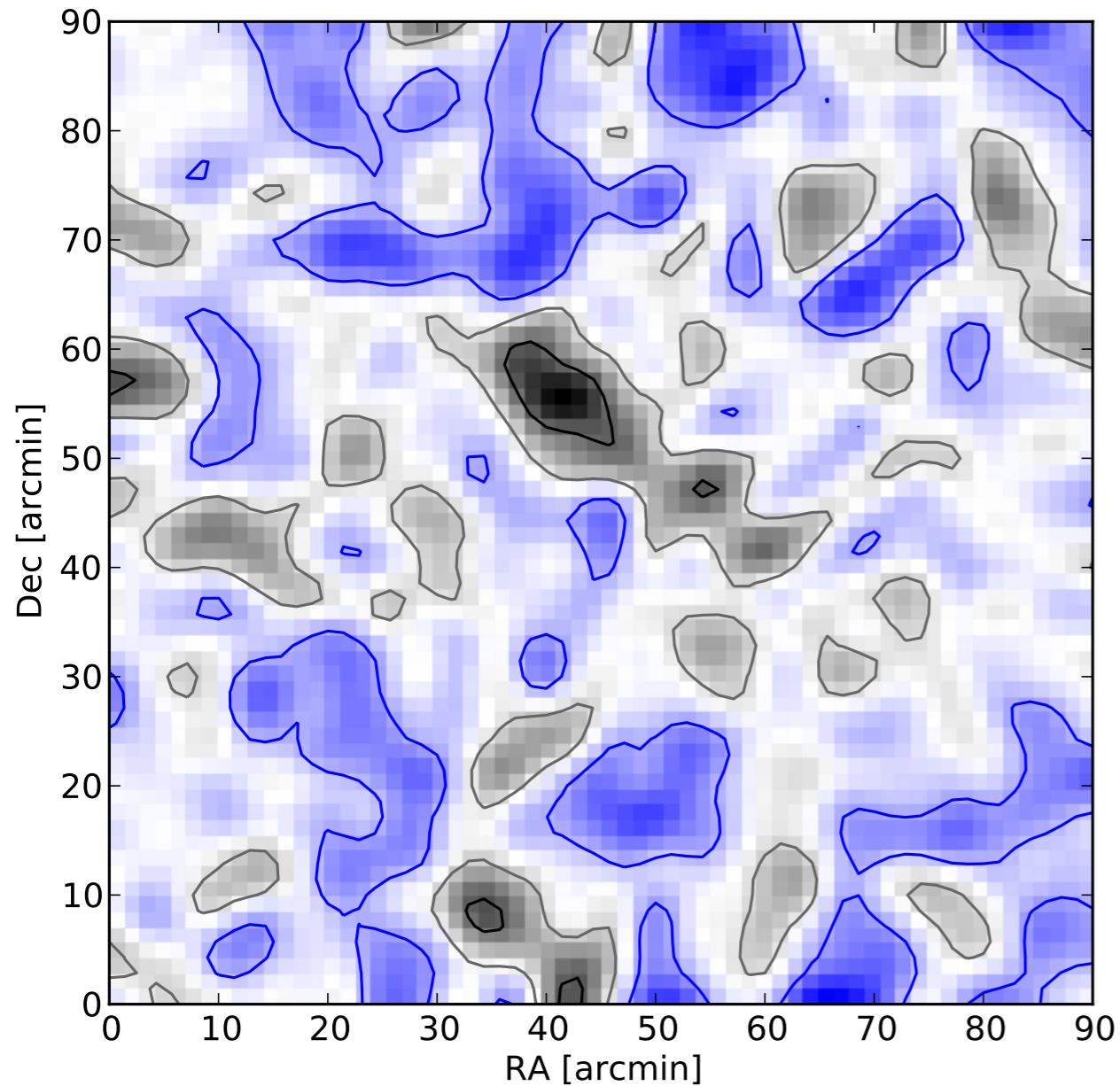
color-cut selection

# Cluster member selection

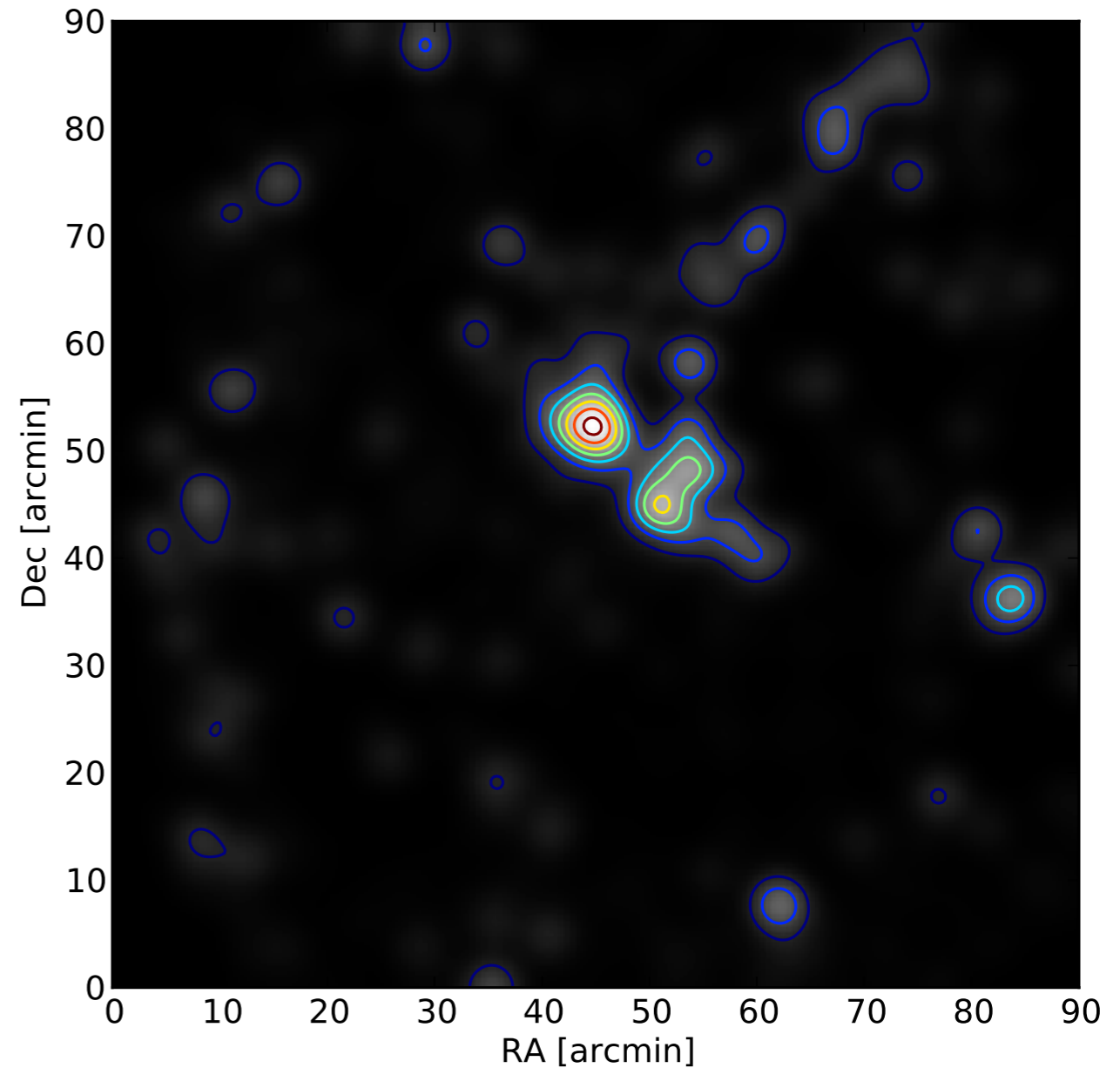
based on color cuts and redMaPPer



# Mass & light maps

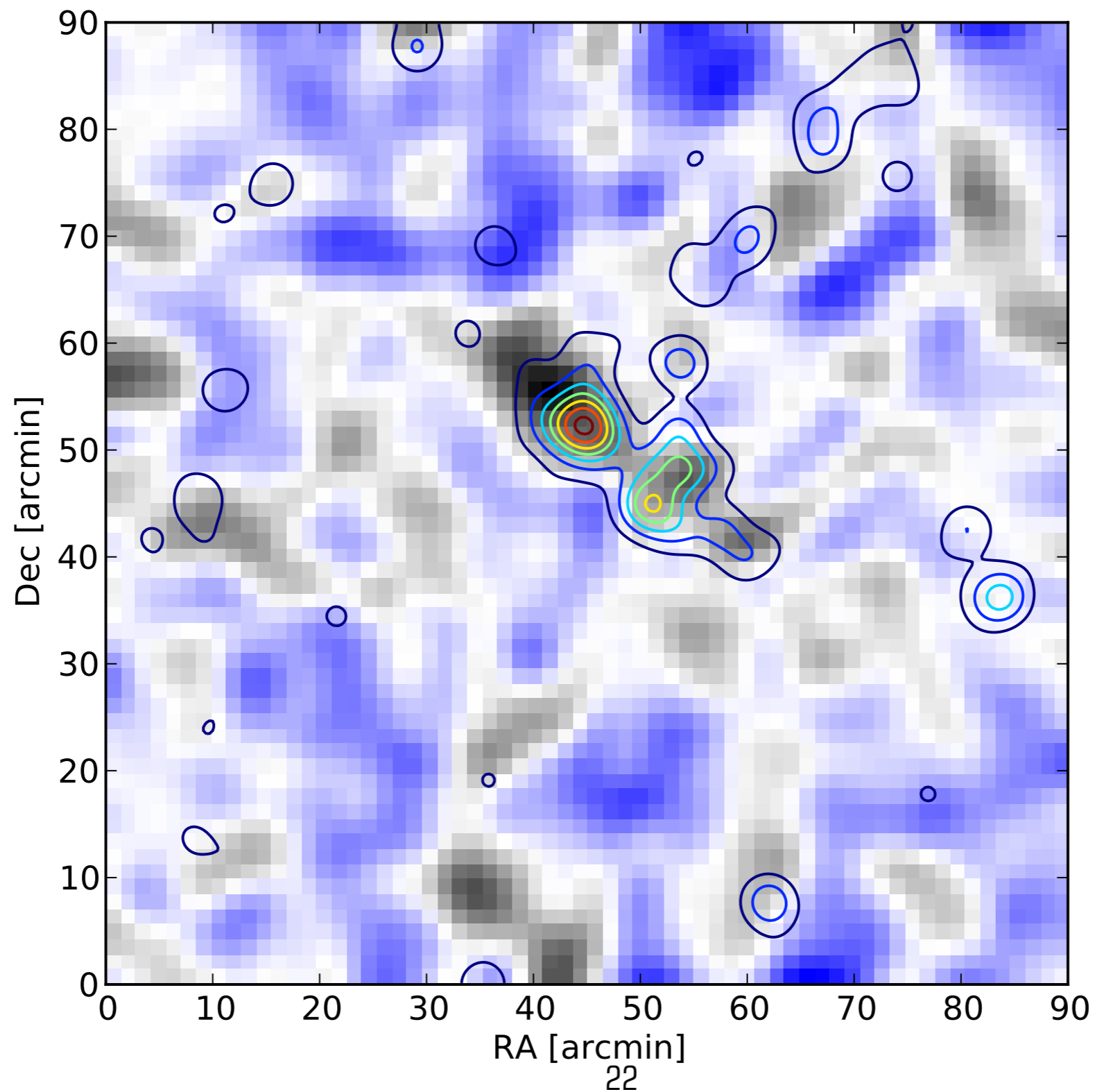


WL mass reconstruction



redMaPPer galaxy distribution at  $z=0.35$

# Mass & light maps

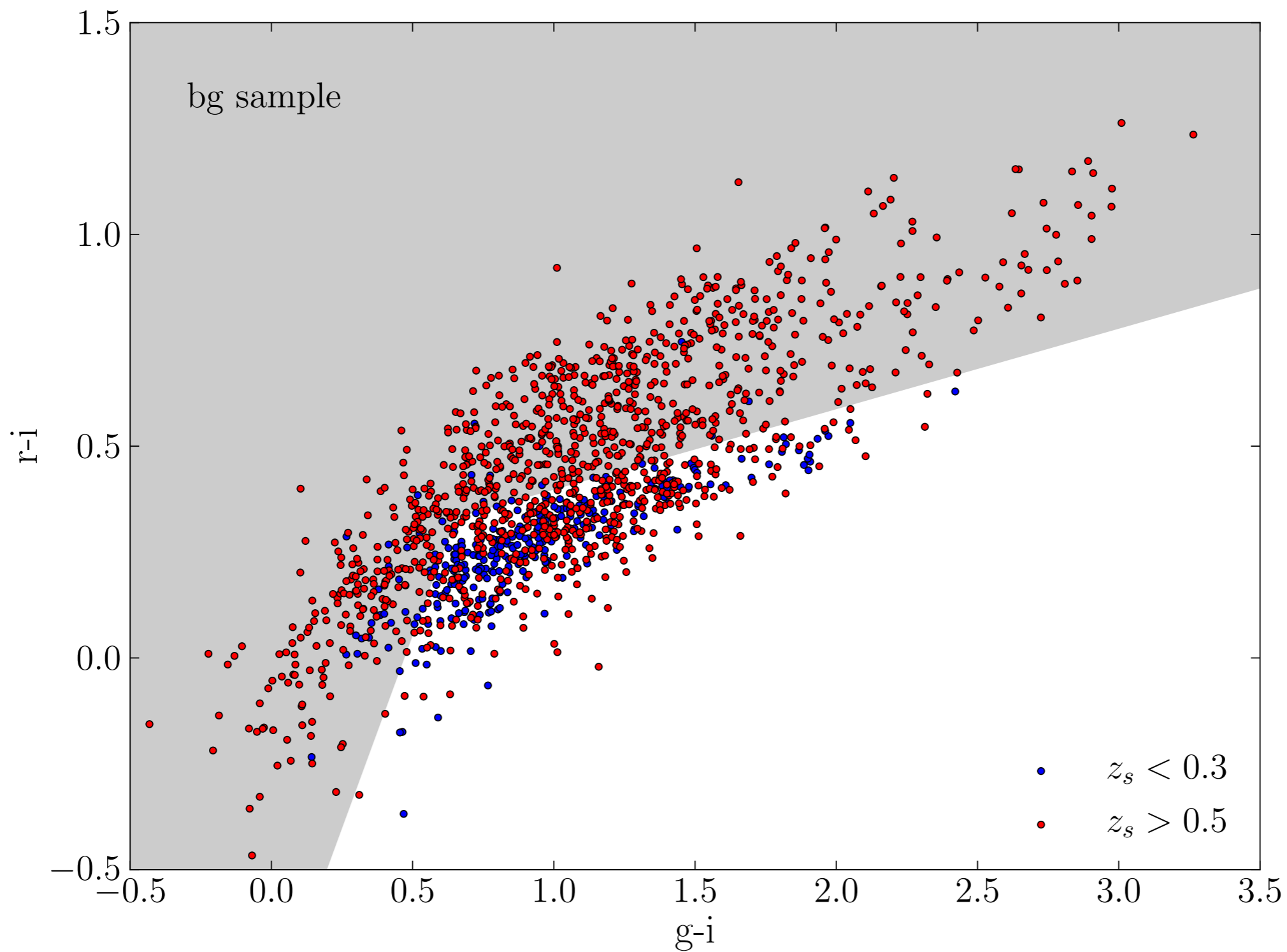


# Outlook



- ❖ Analysis soon completed on 5 cluster fields: 2.3 sq. degrees each
  - ◆  $M_{200} = 4 \cdot 10^{14} \dots 2 \cdot 10^{15} M_{\text{solar}}$
  - ◆ Redshifts  $z = 0.3 \dots 0.9$
- ❖ Investigation on mass-light connection in cluster environments
  
- ❖ Tricky issues
  - ◆ careful treatment of bright cluster galaxies and diffuse cluster light
  - ◆ (faint) cluster member contamination

# Color cuts for background





# Color cuts for background



DARK ENERGY  
SURVEY

